



Femtosecond and nanosecond laser induced breakdown spectroscopic studies of NTO, HMX, and RDX

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ARTICLE INFO

Article history:

Received 5 April 2012

Accepted 29 October 2012

Available online 10 November 2012

Keywords:

Laser induced breakdown spectroscopy

Femtosecond

Nanosecond

High energy materials

CN molecular emission

ABSTRACT

We present our results from the laser induced breakdown spectroscopic studies of 5-Nitro-2,4-dihydro-3H-1,2,4-triazol-3-one (NTO), Octahydro-1,3,5,7-tetranitro-1,3,5,7-tetrazocine (HMX), and 1,3,5-Trinitroperhydro-1,3,5-triazine (RDX) investigated using nanosecond and femtosecond pulses. The presence of C, CN peaks in the spectra, signatures of high energy materials, was confirmed and persistence of emissions has been measured. Some of the Nitrogen peaks in fs LIBS spectra were found to be lower in magnitude (after normalization with N 868.60 nm peak) compared to the ns LIBS spectra. The presence of an additional CN peak in the fs spectra was identified for all samples. The ratio of CN peaks (388.28 nm, 387.08 nm, 386.16 nm) to C peak (247.82 nm), recorded with similar fluences, was discovered to be stronger in the fs case. Some of the possible mechanisms ensuing from our studies towards discrimination of such materials are outlined.

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1. Introduction

Several laser based techniques have been developed recently to investigate the potential of identifying hazardous and illicit compounds [1–5]. The challenges involved in this endeavour are on-field and remote detection to avoid damage to human lives, exposure to harsh environments, fast detection, and negligible scope for false positives. Among the various methods investigated recently, laser induced breakdown spectroscopy (LIBS) has been identified as an attractive, versatile technique with encouraging attributes such as stand-off detection capability, prospective trace material detection, process control, and high speed [6–27]. Continuous efforts have resulted in development of man-portable LIBS and standoff detection successfully tested up to 100 m [28,29]. LIBS has been prolifically evaluated in recent times for detection/identification of explosive residues [15–29], chemical and biological materials [30,31], and landmines [32] etc. Ultrashort laser pulses have tremendous potential in designing novel and sensitive systems for LIBS analysis [33,34]. Conventional LIBS uses nanosecond (ns) pulses whereas short pulse LIBS uses typically femtosecond (fs) pulses. The specific advantages include lower ablation threshold, minimal air entrainment, and reduced background continuum emission, significant attributes required for efficient detection of explosives [35–41]. The LIBS technique has been fairly successful in a variety of applications such as nuclear forensics [42], cultural heritage monitoring [43], identification of bacteria [39–41], etc. to name a few amongst others reported. Although the LIBS technique is

simple, the data analysis is quite cumbersome and occasionally the conclusions drawn are ambiguous. There have been numerous approaches to improve the efficiency such as double/dual pulse LIBS [44,45]. That the LIBS data is sensitive to various parameters including the effects of substrate [46], sample temperature [47], wavelength [48], bandwidth [49], polarization [50], input energy [51] etc. renders it imperative for exploring other mechanisms (e.g. understanding the kinetics/dynamics) towards discrimination/classification of these compounds.

Utilization of ultrashort pulses in the field of ablation, micro-machining, and material processing has been well established [52,53]. In the field of LIBS, Baudelet et al. [39–41] have demonstrated that concentration profile of trace elements enabled them to successfully discriminate different species of bacteria, unambiguously, using fs pulses. Significantly, they demonstrated that (a) fs plasma had lower temperature compared to ns plasma (b) a specific ablation regime that favors intramolecular bonds emission with respect to atomic emission. A precise kinetic study of molecular band intensities allowed them to distinguish the contribution of native CN bonds released by the sample from Carbon recombination with atmospheric Nitrogen. Though there are few initial reports of LIBS experiments with fs pulses on explosive related materials [35–37], the mandate has been divided over advantages of utilizing ultrashort pulses. Earlier studies have attempted to identify a few markers for distinguishing HEM's using fs LIBS data [35–37]. For example, Dikmelik et al. [36] observed emission from CN and C₂ molecules in the LIBS spectra of TNT recorded with 1 mJ, 150 fs pulses. In comparison with fs LIBS the longer timescale of ns excitation resulted in severe fragmentation and, as a consequence, they observed emission only from the elemental constituents in the case of TNT. In contrast to reported fs LIBS spectra of explosives (especially TNT) De Lucia et al. [37] observed

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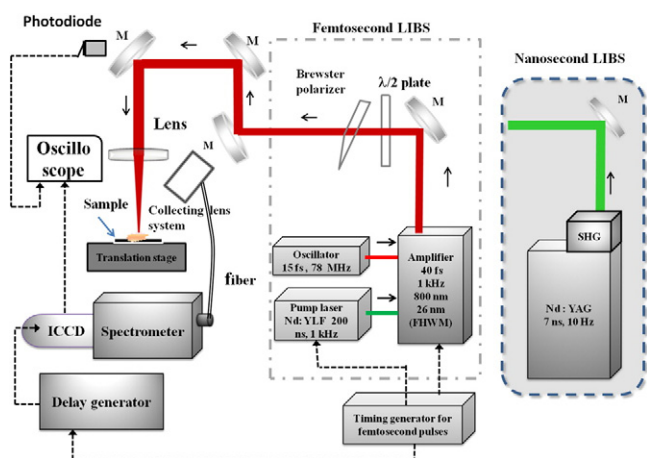


Fig. 1. Schematic of the experimental setup used for ns/fs LIBS.

atomic emission peaks for the constituent elements of explosives – carbon, hydrogen, nitrogen, and oxygen. However, they performed experiments with 10 mJ energy pulses (corresponding to a fluence of 30 J/cm^2). Consequently, further detailed studies are essential to develop infallible methodologies for the identification/classification of HEMs. Herein, we present some of our experimental results from the LIBS measurements of the commonly used high energy materials such as 5-Nitro-2,4-dihydro-3H-1,2,4-triazol-3-one (NTO), Octahydro-1,3,5,7-tetranitro-1,3,5,7-tetrazocine (HMX), and 1,3,5-Trinitroperhydro-1,3,5-triazine (RDX). The LIBS spectra were collected using a gated ICCD in combination with a spectrometer in ambient atmosphere. The persistence of emission of observed atomic and molecular species were evaluated. Detailed analysis revealed that the CN peak magnitude was observed to be stronger in the fs LIBS data compared to ns LIBS data, obtained at similar fluences. The ramifications of these results towards initial stages of discriminating such materials are discussed.

2. Experimental details

Fig. 1 shows the schematic of experimental set up. Nanosecond pulses at 532 nm (10 Hz, 6 ns FWHM, $\sim 600 \text{ mJ}$), amplified fs pulses at

800 nm (1 kHz, $\sim 40 \text{ fs}$ FWHM, $\sim 2.5 \text{ mJ}$) were used for the experiments. The fs amplifier was seeded with $\sim 15 \text{ fs}$ pulses (FWHM, spectral bandwidth of 55–60 nm, FWHM) from an oscillator. That the pulses were nearly transform-limited was confirmed using Silhouette (Coherent, USA) measurements (see supplementary information) based on the MIIPS technique [54]. Other specifications include $\sim 6.5 \text{ mm}$ ($\sim 10 \text{ mm}$) input spot size for ns (fs) pulses and typical energies used were $\sim 20 \text{ mJ}$ ($\sim 1.6 \text{ mJ}$). The laser pulses were focused (using an 80 mm lens) on the sample mounted on a computer controlled XY translation stage. In fs LIBS experiments the laser pulses had a bandwidth of $\sim 26 \text{ nm}$ (FWHM). The samples in both the cases were placed approximately 3–4 mm before the focus (towards the lens) where the spot size was estimated to be $\sim 100 \mu\text{m}$. The estimated peak intensity was $\sim 10^{10} \text{ W/cm}^2$ in the ns case where as in the fs case it was $\sim 10^{13} \text{ W/cm}^2$. The calculated fluences were $\sim 250 \pm 50 \text{ J/cm}^2$ for the ns pulses and in the range of $30 \pm 6 \text{ J/cm}^2$ for the fs pulses. Errors in estimating the fluence arise primarily from evaluation of laser spot size on the sample. Additionally, in order to achieve similar fluences the fs pulses were focused using a 5 cm lens together with slightly higher input energies (2.1 mJ compared to 1.6 mJ in the lower fluence case) providing a fluence of $200 \pm 40 \text{ J/cm}^2$. The light was collected by a lens collection system (UV grade fused silica) and the signal was coupled to a fiber and was transferred to an ICCD + spectrometer system (Andor istar DH734 + Mechelle spectrograph), trigger to the ICCD was provided by a delay generator (SRS, DG 532). Since the detector efficiency was not uniform over the entire spectral range we performed appropriate corrections using standard deuterium and halogen lamps before the data was collected. The samples were prepared in the form of a pellet by first grinding the sample with agate mortar and then applying 10 ton hydraulic pressure onto the powdered sample. The obtained pellets were 2–3 mm thick and were placed on an Al sheet/KBr for recording the spectra. The samples were translated using an XY stage so as to expose a new region for the next laser pulse.

3. Results and discussion

Fig. 2(a)–(c) shows typical LIBS spectra recorded from pellets of NTO, RDX, and HMX, respectively, using ns pulses. The data was recorded with a gate width of 1000 ns and a gate delay of 1000 ns with respect to output of the laser pulse. Before 1000 ns (gate delay) we detected strong Continuum in the ns LIBS data. We observed the

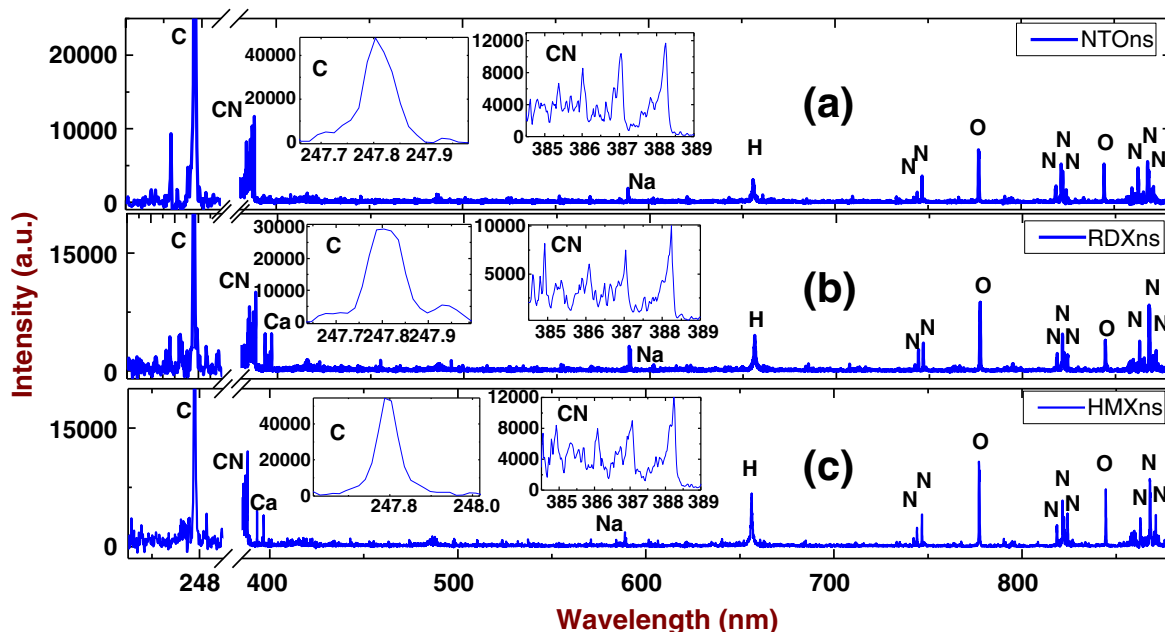


Fig. 2. LIBS spectra of (a) NTO (b) RDX (c) HMX using ns pulses. Gate width of 1000 ns and a gate delay of 1000 ns were used for acquiring the spectra. Each spectrum is resultant of 10 accumulated spectra. Insets show the expanded view of Carbon peak and CN peaks. Fluence used was $250 \pm 50 \text{ J/cm}^2$ with samples placed on Al substrate.

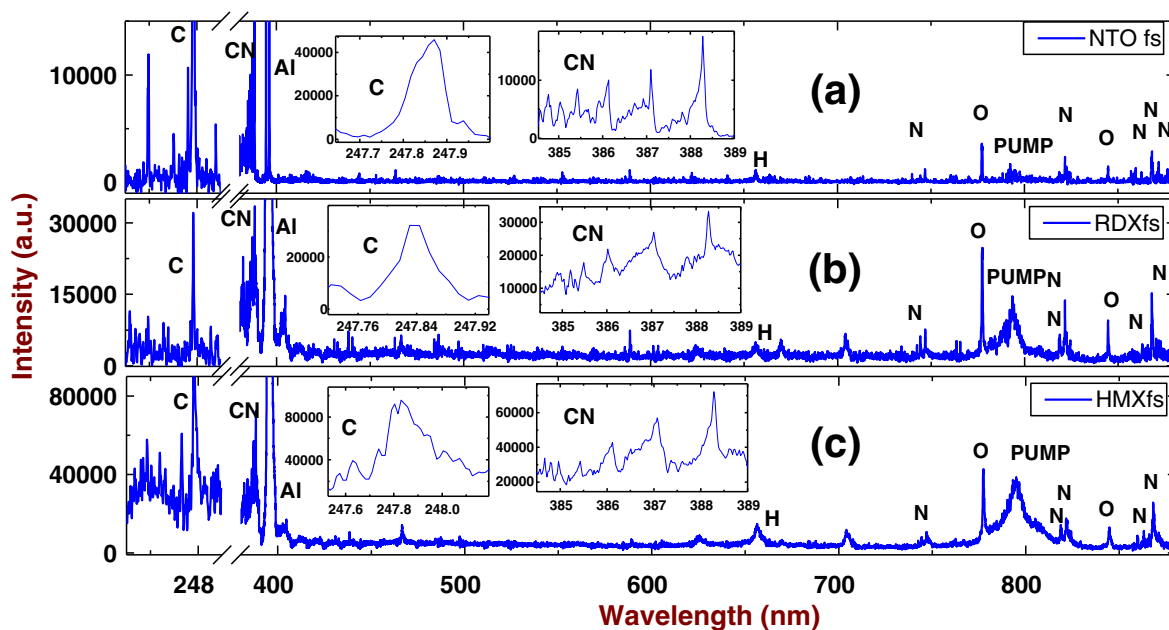


Fig. 3. LIBS spectra of (a) NTO (b) RDX (c) HMX using fs pulses. Gate width used of 300 ns and a gate delay of 100 ns were used for acquiring the spectra. Each spectrum is resultant of 1000 accumulated spectra. Insets show the expanded view of Carbon peak and CN peaks. Fluence used was $30 \pm 6 \text{ J/cm}^2$ with samples placed on Al substrate.

peaks of CN, C, H, N, O in all the spectra. Fig. 3(a)–(c) illustrates the LIBS spectra of NTO, RDX, and HMX, respectively, collected using fs pulses. The data was recorded with a gate width of 300 ns and a gate delay of 100 ns. The data was collected after 100 ns delay to avoid contribution from the continuum. We have also observed aluminum peaks in fs LIBS spectra which were confirmed to be from the substrate since the pulses were focused beyond the sample and on to the substrate. The spectra presented in Figs. 2 and 3 are the result of 10 accumulations in the ns case and 1000 accumulations in the fs case. We noticed that the CN peak at 385.01 nm was clearly present in the fs LIBS spectra while the peak was buried within the noise in the ns spectra. Table 1 summarizes the information of all atomic and molecular peaks observed in fs/ns LIBS spectra of all three molecules studied. Some of the Nitrogen peaks observed in fs LIBS spectra were weaker in magnitude compared to ns spectra (for data examined after normalization with the strongest nitrogen peak at 868.60 nm). Typical magnitudes of nitrogen peaks

Table 1

List of various atomic and molecular peaks in the LIBS spectra of NTO, RDX, and HMX obtained using fs and ns pulses.

Peak	Species	RDX		HMX		NTO	
		ns	fs	ns	fs	ns	fs
247.82	C	*	*	*	*	*	*
279.49, 280.21, 285.10	Mg	*	*	*	*	*	*
385.01	CN	#	*	#	*	#	*
385.40	CN	*	*	*	*	*	*
386.16	CN	*	*	*	*	*	*
387.07	CN	*	*	*	*	*	*
388.28	CN	*	*	*	*	*	*
393.25, 396.77	Ca	*	*	*	*	*	*
394.4, 396.15	Al	*	*	*	*	*	*
588.89, 589.50	Na	*	*	*	*	*	*
656.2	H	*	*	*	*	*	*
742.2, 744.1, 746.8	N	**	*	**	*	**	*
777.20	O (triplet)	*	*	*	*	*	*
818.34, 818.64, 821.50, 822.35,	N	*	*	*	*	*	*
824.22, 859.5, 856.74,	N	**	*	**	*	**	*
862.9, 865.66, 868.03, 870.25,							
871.10, 870.74							
844.55	O	*	*	*	*	*	*
867.80, 868.80	N	*	*	*	*	*	*

* indicates peak present ** indicates higher magnitude # indicates buried in the noise.

evaluated from the raw data were ~2–9 times the noise in fs case whereas they were ~5–22 times the noise in ns case. Differences in the intensities could, probably, be from differences in plasma temperature between the ns and fs case. The Carbon peak at 247.82 nm was observed in both the cases. CN formation in the fs LIBS case would be aided by C from the sample and N from the sample and/or atmosphere.

To understand the temporal dynamics of LIBS signal [55–57] we had recorded spectra with different gate delays. In case of fs LIBS the spectra were recorded with 50 ns gate width starting at 100 ns with 50 ns interval up to 400 ns (total number of 7 spectra were collected) whereas in ns case the spectra were recorded with 200 ns gate width starting at 100 ns with 200 ns interval up to 3000 ns (total number of 15 spectra were collected). The reason we had taken different series size is that in fs LIBS good signal-to-noise (SNR) ratio was observed up to 400 ns only. In the case of ns LIBS we could achieve good SNR up to 3000 ns for all the samples. The intensity of C, H, N, O, and CN peaks recorded with time delay were fitted with a single exponential decay. Decay constants summarized in Table 2 illustrate that the atomic species have longer life time (of few 100's of ns) with ns pulse excitation whereas for fs excitation their decay time was only few 10's of ns. The error in measurements and fittings was estimated to be ~10% and therefore the values presented in Table 2 are within $\pm 10\%$. The decay time of fs LIBS plasma, in general, is very short compared to ns LIBS plasma. The reason is well established and reported/discussed extensively in literature [58–60]. The interaction mechanism of materials is different for these two pulses and, therefore, the nature of plasma is different. In the fs

Table 2

Summary of the persistence of emission of important atomic and molecular species extracted from the LIBS spectra of RDX, HMX, and NTO acquired using ns and fs pulses.

Peak- λ (nm)	Persistence of emission (ns)					
	RDX		HMX		NTO	
	ns case	fs case	ns case	fs case	ns case	fs case
C-247.82	440	68	480	60	510	52
CN-388.28	703	294	812	92	983	133
H-656.20	370	45	379	47	455	36
O-777.20	354	33	218	20	296	29
O-844.55	295	50	379	50	678	37
N-868.80	353	54	300	37	363	30

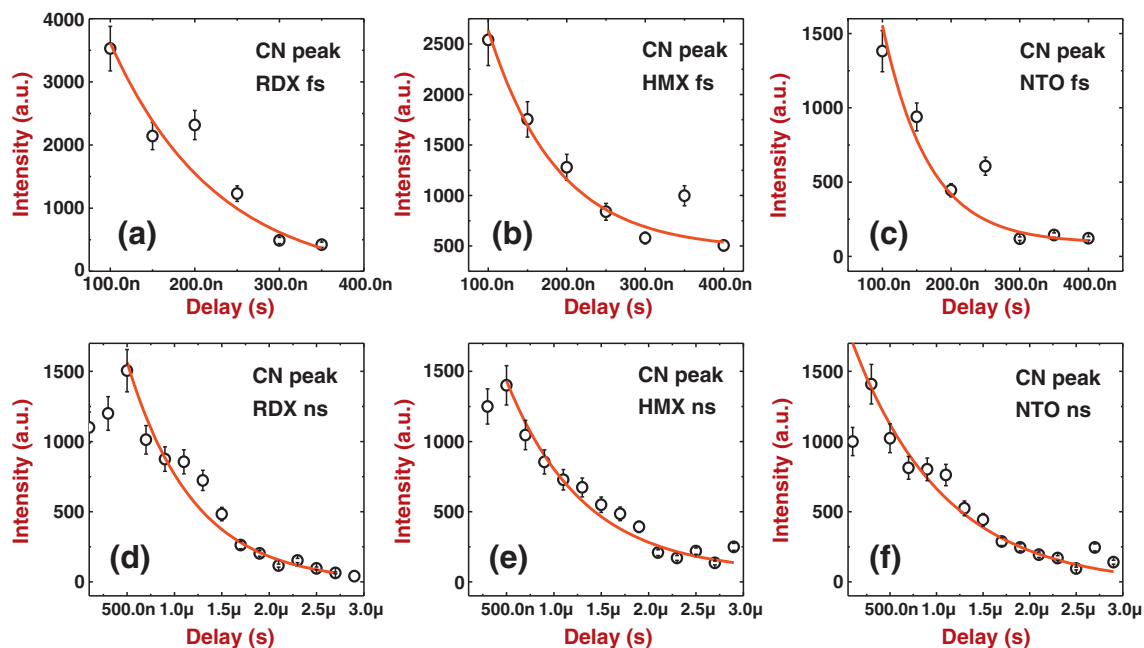


Fig. 4. Time-resolved data of the CN molecular peak (388.28 nm) (a)–(c) fs LIBS data (d)–(f) ns LIBS data recorded with ICCD. 200 ns gate width and an initial gate delay of 100 ns was used for acquiring the spectra and spectra were recorded for every delay of 200 ns up to 2900 ns in the ns case. 50 ns gate width and a starting gate delay of 100 ns was used for acquiring the spectra and spectra were recorded for every delay of 50 ns up to 400 ns in the fs case. Fluence used was 250 ± 50 J/cm² in ns case and 30 ± 6 J/cm² in fs case.

case, material will be rapidly converted into ionized gas. Because of the short time scales involved, the interaction of fs pulses with plasma is absent after certain time whereas ns pulses interact with several transient states of the plasma. Moreover, in the ns case, pulses interact further with the plasma while plasma re-heating takes place which results in a higher temperature and density of the plasma. This leads to a longer persistence of emission in the case of ns plasmas. Fig. 4(a)–(f) illustrates the decay behavior of CN peak (388.34 nm) for all the samples obtained using fs pulses [(a)–(c)] and ns pulses [(d)–(f)]. A closer inspection of the data in the 1–1.5 μs range suggests a small hump, probably

indicating the presence of secondary process. Ma et al. [55] reported their kinetic model of atomic and molecular emission in LIBS studies of organic compounds and demonstrated that four-center reaction $C_2 + N_2 \rightarrow 2CN$ is very unlikely to be the major process during the formation of CN. Rather, the reaction of C and N₂ appeared to be responsible for increased CN concentration at longer delays. Further experiments are required to completely understand this behavior. Fig. 5(a)–(f) illustrates the decay behavior of C peak (247.82 nm) for all the samples obtained using fs pulses [(a)–(c)] and ns pulses [(d)–(f)]. The decay constants are summarized in Table 2. It is evident

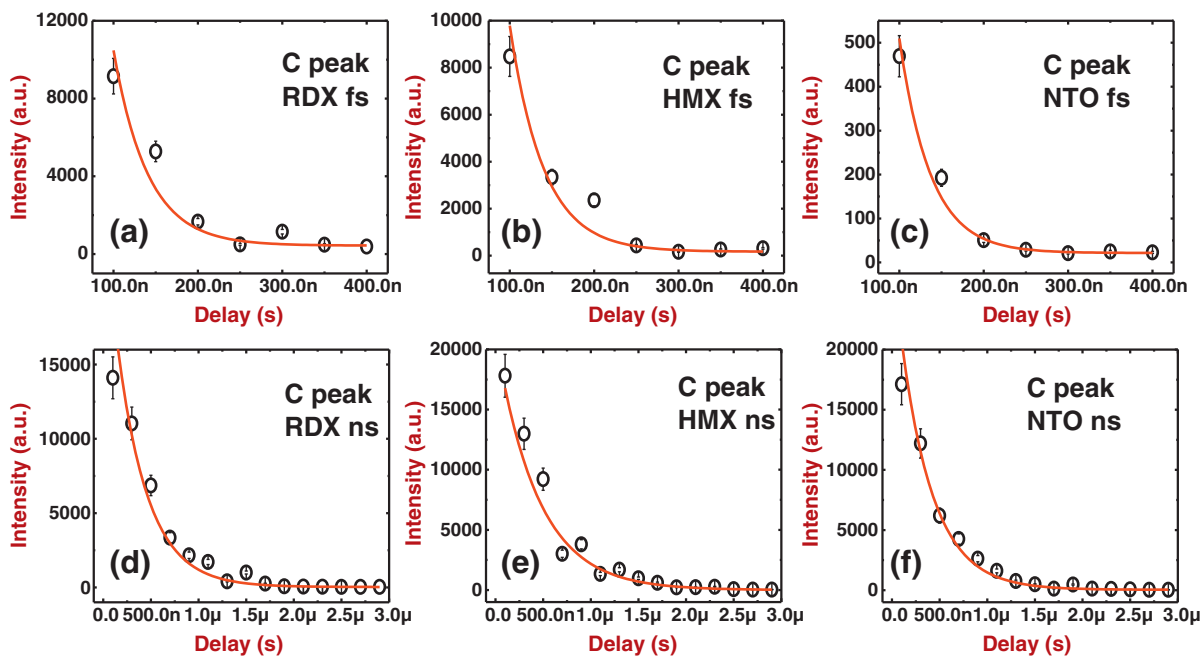


Fig. 5. Time-resolved data of elemental Carbon peak (247.82 nm) (a)–(c) fs LIBS data (d)–(f) ns LIBS data recorded with ICCD. 200 ns gate width and an initial gate delay of 100 ns was used for acquiring the spectra and spectra were recorded for every delay of 200 ns up to 2900 ns in the ns case. 50 ns gate width and a gate delay of 100 ns was used for acquiring the spectra and spectra were recorded for every delay of 50 ns up to 400 ns in the fs case. Fluence used was 250 ± 50 J/cm² in ns case and 30 ± 6 J/cm² in fs case.

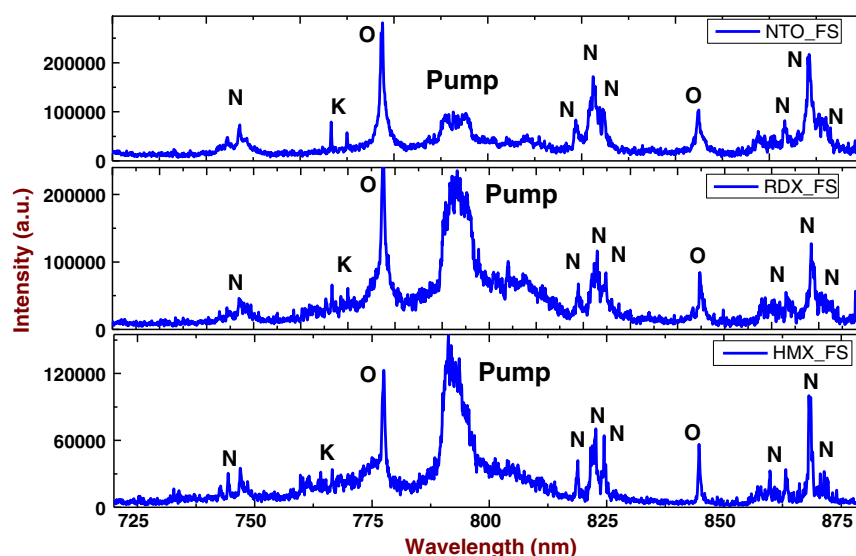


Fig. 6. LIBS spectra of (a) NTO (b) RDX (c) HMX using fs pulses. Gate width used was 800 ns and a gate delay of 100 ns was used for acquiring the spectra. Each spectrum is resultant of 1000 accumulated spectra. Fluence used was $200 \pm 40 \text{ J/cm}^2$ with samples placed on KBr substrate.

that the decay was much slower ($> 400 \text{ ns}$) for the ns C peak compared to the fs C peak (50–70 ns). Further, the decay of CN peak was also slower ($\sim 1 \mu\text{s}$) with ns pulses compared to fs pulses ($< 0.3 \mu\text{s}$). Intensity of all elemental peaks decayed very fast in fs LIBS spectra compared to ns LIBS spectra.

As previously observed [36,39], our data reveals that under current experimental conditions the behavior of molecular species is different for fs excitation compared to ns excitation. The generation of molecular CN species has been identified to be diverse for different compounds [10,61,62]. Three main routes which possibly could contribute to the formation of molecular CN are (a) direct vaporization of the sample containing CN radicals as intermolecular bonds (b) recombination of the C and N from the compound (c) reaction in the plume containing Carbon (from sample) and Nitrogen (from air) [10]. To compare the ratios of molecular to atomic peaks obtained we had recorded fs spectra with a fluence of $200 \pm 40 \text{ J/cm}^2$ and ns spectra with a fluence of $250 \pm 50 \text{ J/cm}^2$ for all the HEMs. The data collected with 800 ns gate width and in the spectral range of 700 nm–900 nm is depicted in Fig. 6(a)–(c) for NTO, RDX, and HMX, respectively. Most of the nitrogen peaks were clearly resolved unlike in the low fluence case, depicted in Fig. 3. Fig. 7(a)–(f) shows the plots CN and C peaks in the LIBS spectra observed in all the samples [(a),(b) for NTO, (c),(d) for RDX and (e),(f) for HMX] using both the pulses. The significant observations from our data were (a) the ratios of CN molecular lines (388.28 nm, 387.08 nm, 386.16 nm, 385.438 nm), to C 247.82 nm line, were stronger in the fs LIBS spectra for all three samples. This is evident from the data presented in Fig. 7 where CN/C ratio is lower in the case ns compared to fs case for all the samples. (b) 385.01 nm peak was clearly observed in fs LIBS spectra while in the ns spectra the peak was buried within the noise. The CN/C ratios estimated for all the samples are summarized in Table 3 and include standard deviation and relative standard deviation (%RSD). For the fs spectra we had calculated the ratios with two different fluences ($30 \pm 6 \text{ J/cm}^2$ and $200 \pm 40 \text{ J/cm}^2$). CN/C ratios were estimated for three strongest CN peaks (388.28 nm, 387.08 nm, and 386.16 nm). The intensity calculated was simply the magnitude of peak height with background correction. The CN (388.28 nm)/C (247.82 nm) ratio estimated for NTO with fs pulses (low/high fluences) was $0.28 \pm 0.08/0.37 \pm 0.09$ while for ns pulses it was 0.25 ± 0.04 . For the case of RDX the ratios were $0.75 \pm 0.06/1.05 \pm 0.40$ and 0.23 ± 0.05 for fs and ns pulses, respectively. The ratios approximated for HMX were $0.61 \pm 0.13/0.82 \pm 0.35$ and 0.19 ± 0.03 for fs and ns pulses, respectively. The CN/C ratios of ns and fs data of all three CN peaks

acquired with similar fluences (and similar gate widths) clearly indicate that the CN formation was indeed superior with fs pulses. The typical signal to noise ratio was > 20 in the ns case and > 25 in the fs case. Similar trend was observed with LIBS studies on bacteria samples in which they report CN intensity was high in fs LIBS spectra compared to the ns LIBS spectra [39]. These ratios provide useful insight during the discrimination of these compounds. For example, one could look at (a) the presence of C, CN, C_2 peaks (b) CN/C, CN/ C_2 and other molecular and atomic ratios as performed successfully by Gottfried et al. [27] (c) the time evolution of CN peak to eliminate Nitrogen based molecules etc.

De Lucia et al. [37] have investigated the efficacy of fs LIBS for detection of explosive residues of RDX, C-4, and Composition-B. They observed a lower threshold for LIBS experiments with $\sim 120 \text{ fs}$ pulses in contrast to ns pulses. In RDX spectra they observed stronger C (near 248 nm) emission with fs pulses compared to ns pulses acquired with similar fluences. They have demonstrated that any advantages of fs LIBS over ns LIBS are only likely to be realized at the very low fluences possible with fs pulses. They utilized a gate delay of 50 ns with pulses of 10 mJ energy (corresponding to a fluence of 30 J/cm^2) with a gate width of 5/10 μs and observed Oxygen and Nitrogen emission from atmospheric entrainment in the blank Al spectrum. Our data obtained with $\sim 40 \text{ fs}$ pulses and lower fluences (see supplementary information) with pure Al substrates clearly suggested minimal air entrainment. However, the data recorded at higher fluences using fs pulses avoided Al substrate since the samples were placed on a KBr substrate. The samples used were also thicker ($\sim 3 \text{ mm}$) to further reduce air entrainment. Most importantly, our pulses have extremely large bandwidth associated with them ($\sim 28 \text{ nm}$ FWHM corresponding to a total bandwidth of $\sim 65 \text{ nm}$ implying that all wavelengths in the $800 \pm 32 \text{ nm}$ spectral range were available). There is a strong possibility that some of these wavelengths could be resonant with the CN transition (in the spectral range of 384 nm–389 nm) via multi-photon absorption (e.g. 2-photon absorption in this case) thereby inducing slightly larger emission. With the ns pulses the bandwidth involved is extremely small ($< 1 \text{ nm}$) consequently lacks any resonance. That there was no peaking of the CN intensity with gate delay [63,64] in the fs LIBS data also indicates the contribution is mainly from the constituent C/N or CN of the samples. The plasma chemistry is rather complex in such samples with the possibility of Carbon reacting with Oxygen to form CO (followed by CO_2) and therefore, in the ns case, CN formation could be partially inhibited [56,65,66]. Experiments with exact fluences and devoid of atmospheric intervention are also necessary to absolutely identify the magnitude of

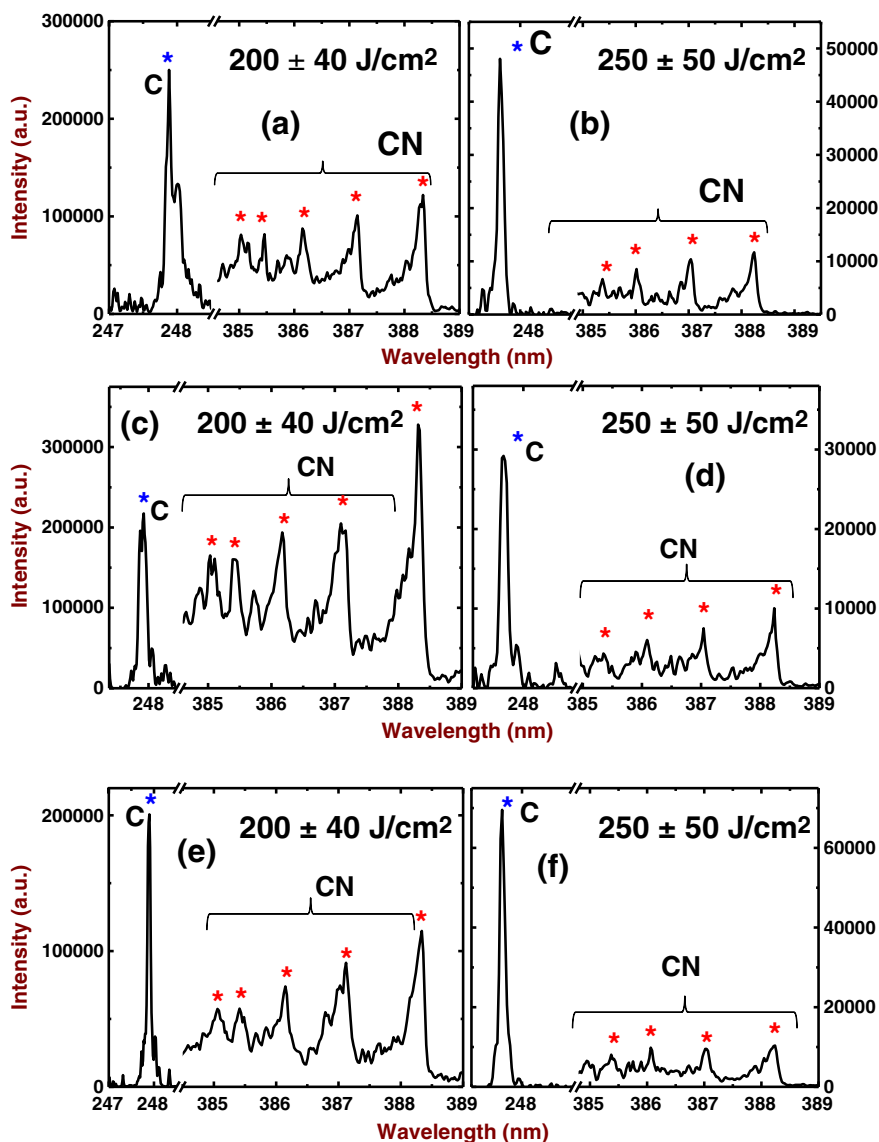


Fig. 7. Comparison of the C and CN peaks obtained in fs, ns LIBS spectra of (a), (b) NTO (c), (d) RDX and (e), (f) HMX. Gate width of 1000 ns and a gate delay of 1000 ns were used for acquiring the ns spectra. Gate width used of 800 ns and a gate delay of 100 ns were used for acquiring the spectra. *s are guide to peak positions.

CN/C ratios and our efforts are directed towards achieving these. We are in the process of developing appropriate methodologies towards discrimination of HEMs similar to those successfully implemented [67,68]. The present work provides us with an idea of the nature of different pulses (ns/fs) interacting with these materials. Moreover, the behavior of spectral features varying with pulse duration is also understood.

4. Conclusions

We have studied the LIBS spectra of NTO, RDX, and HMX using ns and fs pulses recorded in ambient atmosphere. The elemental and molecular features observed in both spectra were compared. The presence of C and CN peaks in LIBS spectra, signatures for the presence of high energy materials, was confirmed and their persistence

Table 3

CN/C ratios evaluated for NTO, RDX, and HMX for three strongest peaks of CN. The data is an average obtained from 6 independent measurements in each case.

Sample	Ratio	fs – 30 ± 6 J/cm ²	fs – 200 ± 40 J/cm ²	ns – 250 ± 50 J/cm ²
NTO	CN-388.28/C-247.82	0.28 ± 0.08(26.63)	0.37 ± 0.09(24.71)	0.25 ± 0.04 (17.7)
	CN-387.08/C-247.82	0.21 ± 0.06(28.08)	0.28 ± 0.06(22.17)	0.20 ± 0.04 (19.8)
	CN-386.16/C-247.82	0.19 ± 0.04(19.52)	0.22 ± 0.06(25.30)	0.17 ± 0.06(33.41)
RDX	CN-388.28/C-247.82	0.75 ± 0.06(8.15)	1.05 ± 0.40(37.67)	0.23 ± 0.05 (20.2)
	CN-387.08/C-247.82	0.53 ± 0.06(11.20)	0.79 ± 0.28(35.71)	0.20 ± 0.06(29.42)
	CN-386.16/C-247.82	0.37 ± 0.06(16.01)	0.63 ± 0.24(38.42)	0.18 ± 0.10(53.74)
HMX	CN-388.28/C-247.82	0.61 ± 0.13(21.59)	0.82 ± 0.35(45.38)	0.19 ± 0.03(17.76)
	CN-387.08/C-247.82	0.42 ± 0.09(22.41)	0.58 ± 0.24(43.71)	0.15 ± 0.03(19.54)
	CN-386.16/C-247.82	0.25 ± 0.10(40.44)	0.45 ± 0.17(39.81)	0.12 ± 0.03(21.09)

Values are in the format – “Mean ± SD (%RSD).”

of emission has been measured. The persistence of emission of atomic and molecule peaks C, H, N, O, and CN were typically few tens of ns in the fs case and few hundreds of ns in the ns case. One of the CN peaks (at 385.01 nm) was clearly observed in fs LIBS spectra while in the ns spectra the peak was buried within the noise. From the LIBS spectra recorded with similar fluences we observed that the magnitude of CN/C was stronger in the fs case compared to ns case. Further studies in the direction of correlation analysis could provide us a better route for classification of these compounds.

Acknowledgments

We are grateful to the referees for their invaluable comments and suggestions. Financial support from DRDO, India is gratefully acknowledged. The authors also acknowledge the samples provided by HEMRL, Pune, India and P. Ravi, ACRHEM.

Appendix A. Supplementary data

Supplementary data to this article can be found online at <http://dx.doi.org/10.1016/j.sab.2012.11.002>.

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